Model Data Fusion: developing Bayesian inversion to constrain equilibrium and mode structure


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Recently, a new probabilistic “data fusion” framework based on Bayesian principles has been developed on JET and W7-AS. The Bayesian analysis framework folds in uncertainties and inter-dependencies in the diagnostic data and signal forward-models, together with prior knowledge of the state of the plasma, to yield predictions of internal magnetic structure. A feature of the framework, known as MINERVA (J. Svensson, A. Werner, Plasma Physics and Controlled Fusion 50, 085022, 2008), is the inference of magnetic flux surfaces without the use of a force balance model. We discuss results from a new project to develop Bayesian inversion tools that aim to (1) distinguish between competing equilibrium theories, which capture different physics, using the MAST spherical tokamak; and (2) test the predictions of MHD theory, particularly mode structure, using the H-1 Heliac.

Keywords: Bayesian inference, equilibrium, tokamak, stellarator

1. Introduction

With the advent of large scale neutral beam heating, magnetically confined fusion plasmas have drifted from the simple picture of ideal magnetohydrodynamics (MHD), which describes the plasma as a single, stationary, isotropic Maxwellian fluid. Due principally to neutral beam heating, several tokamak experiments now boast plasma toroidal rotation speeds that approach the thermal Mach speed and have significant stored energy residing in the energetic particle population produced by charge exchange with thermals. Motivated by these developments, a range of new descriptions have emerged that include thermal rotation [1] as well as energetic particles [2]. Despite this, ideal MHD is still the foundation of nearly all analysis. Detailed magnetic reconstruction based on this treatment ignores the energetic complexity of the plasma, and can result in model-data inconsistencies, such as thermal pressure profiles which are inconsistent with the total stored kinetic energy of the plasma.

A parallel development has been the improvement in the diversity, accuracy and resolution of plasma diagnostics. Interpretation, however, often requires a detailed knowledge of the plasma equilibrium. For example, inference of the toroidal current profile \(j_\theta(\psi)\) from line of sight measurements of the polarization angle requires a knowledge of the poloidal flux \(\psi\) across the plasma. Formally, diagnostic forward functions relate the vector of plasma parameters \(I\) to the measurement vector \(D\). For a linear system, such as toroidal current inference in a double null configuration, \(I\) and \(D\) are normally related through a response matrix \(M\) with additional contributions \(C\), such that \(D = MI + C\). Inference involves inverting this relationship to give plasma parameters \(I\) that are consistent with the data \(D\). A widespread technique used is least-square fitting, in which prior assumptions are included via a penalty term in the fit.

The confluence of higher performance plasmas with diagnostic development has however led to a dichotomy: data is often inconsistent with ideal MHD equilibria, sometimes misleading scientists to propose new phenomena to describe data “artifacts”. An example is filamentation of flux surfaces in the Rijnhuizen Tokamak, which was inferred from fluctuations in measurements of electron temperature [3]. Subsequent analysis of the Thomson scattering detection chain revealed that with the correct photo-electron statistics and 2D instrument profile, similar structures could arise from noise. [4, 5].

Recently, a new integrated data-modelling approach for inference of fusion plasma parameters has emerged which offers a natural framework with which to resolve artifacts from model behaviour. In contrast to least square fitting, the Bayesian approach to inference in fusion plasmas, developed by multiple authors, [6, 7, 8, 9, 10, 11, 12, 13] involves the specification of an initial prior probability distribution function (pdf), \(P(I)\), which is then updated by taking into account information that the measurements provide through the likelihood pdf \(P(D|I)\). The result is the posterior distribution \(P(I|D)\) given by Bayes’ formula

\[
P(I|D) = \frac{P(D|I)P(I)\,dI}{P(D)}.
\]

The advantage of the Bayesian approach over traditional inversion techniques is two-fold: (i) prior knowledge, including known parameter inter-dependencies is made explicit, and (ii) as the formulation is probabilistic, random errors, systematic uncertainties and instrumental bias are integral part of the analysis rather than an afterthought.

In this work we present initial results of forward mod-
els of current tomography in the Mega Ampere Spherical Tokamak (MAST), we formulate Bayesian inference of force balance, and we identify plans for inference of mode structure in the H1 heliac. The paper is structured as follows: Sec. 2 briefly outlines the MAST experiment and present first results for Bayesian inference of poloidal flux and current profiles from MAST motional Stark effect (MSE) measurements. Section 3 outlines a Bayesian inference framework for force balance, synthesizes a pressure profile posterior given the flux surfaces from MSE and Thomson Scattering data, and computes an estimate of \( f(\psi) f^*(\psi) \), with \( f(\psi) \) the toroidal flux function. Using this, we are able to quantify the impact of poloidal currents on the equilibrium configuration. Section 4 outlines a Bayesian inference model for mode structure and develops a plasma model to compute candidate modes in the H-1 heliac. Finally, Sec. 5 contains concluding remarks.

2. MAST and current tomography

The Mega Ampere Spherical Tokamak, which first published physics results in 2001 [14], is one of the world’s largest spherical tokamaks. While MAST physics and technology development have contributed across a broad range of fusion science [15, 16], two properties of MAST help motivate this research: high performance, and precision diagnostics. It is the combination of these properties in a spherical tokamak, which has a relatively weak toroidal field and therefore large sensitivity to poloidal currents compared to conventional tokamaks, which helps motivate selection of MAST to develop Bayesian inference for force balance.

Recently, both MAST neutral beam injectors have been upgraded to 3.8 MW. This has enabled plasma performance to be routinely lifted above \( \beta_n \approx 5 \), which was reported in 2005 [17]. MAST is equipped with an array of precision diagnostics [16], including a high spatial resolution, single-time point, ruby Thomson Scattering system, a multi-time point Nd:YAG Thomson Scattering system with a coarser spatial resolution, Motional Stark effect, Charge Exchange Recombination, and fast magnetics [18].

Our development of Bayesian inference of the current profile on MAST closely follows the seminal work of Svensson and Werner [11]. In that work, the plasma was represented as a grid of toroidal axisymmetric current beams, each with rectangular cross-section and each beam carrying a uniform current density. In MAST, we have placed these beams so as to fill-out the entire limiter region as depicted in Figure 1. The magnetic field generated is a summation of Biot-Savart’s law over current beams.

A key advantage to using a series of current beams with finite cross-section to model the plasma current (as opposed to a filamentary model) is that the semi-analytic expressions for the corresponding magnetic field and vector potential have no singularities, even at points within the current beam itself. Indeed, if one were to use filaments to model the plasma current, there would be many singular points in the calculated magnetic field within the plasma that would make subsequent flux-surface calculations difficult and somewhat questionable [11].

2.1 Diagnostic Signals

In axisymmetric devices, there exists a direct relation between the poloidal flux, \( \psi \), and the toroidal component of the current vector. Specifically, defining the cylindrical coordinate system \((R, Z, \phi)\) with \( R \) the major axis, \( Z \) the vertical axis and \( \phi \) the geometric toroidal angle, the following relation holds:

\[
\psi(R, Z) = \int B \cdot dS = \oint A \cdot dl = 2\pi RA_\phi.
\]
where \( A_\phi \) is related to the toroidal current \( j_\phi \) via Biot-Savart’s law:
\[
A_\phi (r) = \frac{\mu_0}{4\pi} \int \frac{j_\phi (r')}{|r - r'|} dV',
\]
where \( \mu_0 \) is the permeability of free-space. In Eq. (3) \( r \) is the position vector \( r = R \hat{r} + Z \hat{Z} + \phi \hat{\phi} \), which \( e_R, e_Z \), and \( e_\phi \) unite vectors in the \( R, Z \) and \( \phi \) directions, respectively. By inspection, the \( A_R \) and \( A_Z \) components due to \( j_\phi \) are zero. Taking the curl of Eq. (2) and utilizing \( \mathbf{B} = \nabla \times \mathbf{A} \) it can be shown
\[
B_R = \frac{1}{R} \frac{\partial \psi}{\partial Z}, \quad B_Z = \frac{1}{R} \frac{\partial \psi}{\partial R},
\]
where \( B_R, B_Z \) and \( A_\phi \) are the components of the magnetic field. From Eqs. (2)–(5), it can be seen that if \( j_\phi \) is represented by an axisymmetric beam of rectangular cross section and uniform current distribution, then \( \psi, B_R \) and \( B_Z \) will vary linearly with respect to the total current going through that beam. The importance of this will become clear in the next two paragraphs.

The main diagnostics used to infer the current distribution are pickup coils \( P_i \), full flux loops \( F_i \) and the polarisation angle \( \gamma_i \) of the emitted light from neutrally excited species during neutral beam injection due to the motional Stark effect (MSE). The responses of these diagnostics to a current running through a beam, \( I \), is given by
\[
P_i(R, Z; I) = B_R(R, Z; I) \cos(\theta_i) + B_Z(R, Z; I) \sin(\theta_i),
\]
\[
F_i(R, Z; I) = \psi(R, Z; I),
\]
\[
\tan \gamma_i(R, Z; I) = \frac{A_0 B_Z(R, Z; I) + A_1 B_R(R, Z; I) + A_2 B_\phi(R, Z; I)}{A_3 B_Z(R, Z; I) + A_4 B_R(R, Z; I) + A_5 B_\phi(R, Z; I)},
\]
where \( \theta_i \) is the angle between a pickup coil’s normal and the outboard midplane, and \( A_0, A_1, A_2, A_3, A_4 \), and \( A_5 \) are constants for the particular MSE viewing geometry. The notation “…” denotes that the subsequent argument, in this case current \( I \), is held constant. (By convention, the notation “…” is reserved for probability theory.) Both \( P_i \) and \( F_i \) have a linear dependence on \( \psi, B_R \) and \( B_Z \) and hence they are linearly dependent on the total current going through each beam of the current plasma beam model. The function \( \tan \gamma_i \) has a non-linear dependence on \( B_R \) and \( B_Z \); however given that both neutral beams and MSE viewing optics all are on the mid-plane for MAST, \( A_1, A_2, A_3, A_4 \) are all approximately zero. The following approximation is hence used for \( \tan \gamma_i \):
\[
\tan \gamma_i(R, Z; I) \approx \frac{A_0 B_Z(R, Z; I) + A_1 B_R(R, Z; I)}{A_3 B_Z(R, Z; I)},
\]
where we have retained the term \( A_1 \). Since the current beam plasma model does not take into account poloidal currents, vacuum field values for \( B_\phi \) are used in Eq. (9) to perform the current tomography calculations. The correction to \( B_\phi \) due to poloidal currents is investigated in Sec. 3.1.

Thus, given that \( P_i(R, Z; I), F_i(R, Z; I) \) and \( \tan \gamma_i(R, Z; I) \) all have a linear dependence on the current flowing through \( I \), one may write out a generalized prediction vector \( \mathbf{P} \) for all the pickup coils, flux loops and polarization angles, as
\[
\mathbf{P} = \mathbf{M} \mathbf{I} + \mathbf{C}.
\]

Here, \( \mathbf{M} \) is the response matrix of the current vector \( \mathbf{I} \) corresponding to all the plasma beams modelling the plasma, and \( \mathbf{C} \) represents various other contributions to the prediction vector, which are constant relative to \( \mathbf{I} \).

### 2.2 CAR Prior and Bayesian Inference

In [11], Svensson and Werner used a Conditional Auto-regressive (CAR) prior to perform Bayesian inference for the current beam model of the plasma, and this is the model adopted in this paper. The advantage of this choice in prior is that plasma beam currents are correlated to each other but in a spatially localised way. Thus, this prior has the effect of enforcing some smoothness between adjacent current beams, while minimising spatially long ranged effects due to manipulating the current in a particular beam. This is desirable in that such behaviour allows one to get a clear relation between individual current beams and diagnostic measurements. To construct this prior distribution, the following distribution over all current beams is considered
\[
\frac{P[I]}{\exp \left( -\frac{1}{2} \mathbf{I}^T \mathbf{Q} \mathbf{I} \right)},
\]
where the superscript \( T \) denotes the transpose, and \( \mathbf{Q} \) is the precision matrix. Equation (11) is proportional to a zero-mean normal distribution, and so the conditional distribution of one current \( I_i \), given all others \( I_{-i} \) satisfies
\[
p(I_i|I_{-i}) \propto \exp \left( \frac{1}{2} \sum_i \beta_i I_i^2 / (2\tau_i) \right),
\]
for some \( \beta_i \),. Matching terms between Eq. (11) and Eq. (12) shows that \( \beta_i = 0, \beta_{ij} = -Q_{ij}/Q_{ii} \) and \( \tau_i = 1/Q_{ii} \). Moreover, it turns out the symmetry of \( \mathbf{Q} \) implies that \( \beta_{ij} \tau_j = \beta_{ji} \tau_i \) [19]. We have set all variances equal. By setting \( \beta_i \) to correspond to \( I_i \) having a mean that is simply the average of the currents of the beams horizontally and vertically adjacent to it, the precision matrix takes the form
\[
\mathbf{Q} = \frac{1}{\tau} \left( 1 - \frac{1}{4} \mathbf{W} \right),
\]
where \( I \) is the identity matrix and \( \mathbf{W} \) the adjacency matrix with \( W_{ij} = 1 \) if \( i \) and \( j \) are horizontally or vertically adjacent to current element \( I_i \), and \( W_{ij} = 0 \) otherwise.

From Eq. (12), it can be seen that the CAR prior is constructed by exactly specifying the conditional distribution \( p(I_i|I_{-i}) \) and subsequently transforming this expression into a zero-mean normal distribution for \( I_i \). CAR distributions are discussed in detail in [19] and have the advantage...
that hidden parameter inter-dependencies associated with direct manipulation of the covariance matrix in a normal distribution are totally avoided in the CAR construction (see [11] for a discussion on this point).

We have used a normal distribution for every diagnostic measurement \( D \) as it is the least informative, or maximum entropy distribution when only the standard deviation (“error”) and mean value is available for each diagnostic signal [20], as is the case for MAST diagnostic data signals. Under this assumption, the likelihood \( P(D|I) \) becomes a multivariate normal distribution of the data \( D \), and is given explicitly by

\[
p(D|I) = \frac{1}{(2\pi)^{N_D}/2 |\Sigma_D|^{1/2}} \exp \left[-\frac{1}{2} (P - D)^T \Sigma^{-1} (P - D) \right],
\]

where \( N_D \) is the number of measurements, and \( \Sigma \) is the covariance matrix of the measurements, which is determined experimentally. Finally, substituting Eq. (11) and Eq. (14) into Bayes’ formula shows that the posterior distribution \( p(I|D) \) satisfies the following relation:

\[
p(I|D) \propto \exp \left[-(P - D)^T \Sigma^{-1} (P - D) - I^T Q I \right].
\]

Recalling Eq. (10), it is known that \( P \) is linearly dependent on \( I \); and thus, it can be shown that Eq. (15) is proportional to a multivariate normal distribution in \( I \) (see [11] for details). Hence, \( p(I|D) \) has a simple analytic representation that be directly analysed without resorting to using Markov-chain Monte-Carlo (MCMC) algorithms. The speed afforded by the analytic nature of the posterior makes this analysis amenable to real-time plasma control applications.

With the maximum and variance of \( p(I|D) \) determined, Eqs. (2) and (3) can be used to construct the poloidal flux surfaces corresponding to the contours of \( \psi(R,Z) \). Figure 2 shows poloidal flux surfaces from MAST discharge # 22254 at 320 ms using pickup coils, flux loops and MSE. Discharge #22254 is a deuterium plasma in a double-null configuration, which was heated with 3.1 MW and MSE. Discharge # 22254 at 320 ms using pickup coils, flux loops (\( \psi \) poloidal flux surfaces corresponding to the contours of \( Z \)). Figure 2 shows poloidal flux surfaces from MAST shot # 22254 at 320 ms using pickup coils, flux loops and MSE. Figure 2: Poloidal flux surfaces inferred for MAST shot # 22254 at 320 ms using pickup coils, flux loops and MSE.

### 3. Inference of force balance

The physics goal of our work is to exploit the improved resolution of diagnostics to infer the validity of different force balance descriptions. Specifically, we have in mind development of a framework that will validate energetic plasma resolved force balance models. To build towards this goal, we commence by applying Bayesian inference to the axisymmetric Grad-Shafranov equilibrium description of ideal MHD force balance, \( \mathbf{J} \times \mathbf{B} = \nabla P \).

The well known Grad-Shafranov equation [21] can be written

\[
F(R,Z) = 0
\]

where

\[
F(R,Z) = -\mu_0 R j_\phi(\psi) + \mu_0 R^2 p'(\psi) + \mu_0^2 f(\psi) f'(\psi),
\]

\[
-\mu_0 R j_\phi(\psi) = R \frac{\partial}{\partial R} \left( \frac{\partial}{\partial R} + \frac{\partial^2}{\partial Z^2} \right).
\]

For a real plasma, the presence of nonideal effects will mean \( F(R,Z) \) is nonzero. Our long term aim is to compute \( P(\mathcal{F}|D) \) using Bayes formula, such that \( P(D|\mathcal{F}) = P(D|\mathcal{F}) P(\mathcal{F}) / P(D) \). Here, the vector \( D \) comprises all the...
magnetics, motional Stark effect, Thomson scattering and charge exchange recombination data, while the column vector $\mathcal{F}$ contains $F(R, Z)$ evaluated from Eq. (17) at different $(R, Z)$ across the plasma.

An integral representation for the prior $P(\mathcal{F})$ can be obtained by using the transformation property

$$P(\mathcal{F}, p'(\psi), f(\psi)f'(\psi)) \times \frac{\partial(\mathcal{F}, p'(\psi), f(\psi)f'(\psi))}{\partial(j_\phi(\psi), p'(\psi), f(\psi)f'(\psi))} = P(j_\phi(\psi), p'(\psi), f(\psi)f'(\psi))$$

and integrating both sides across $p$ and $f$. That is,

$$P(\mathcal{F}) = \int \frac{P(j_\phi(\psi), p'(\psi), f(\psi)f'(\psi))}{\frac{\partial(\mathcal{F}, p'(\psi), f(\psi)f'(\psi))}{\partial(j_\phi(\psi), p'(\psi), f(\psi)f'(\psi))}} dp'df'$$

where the integrand $dp'df' = dp'(\psi)df(\psi)f'(\psi)$. An ideal MHD plasma satisfies Eq. (16), such that the residual force is zero, and so $P(\mathcal{F}) = \delta(\mathcal{F})$. Unfortunately, the joint distribution function $P(j_\phi(\psi), p'(\psi), f(\psi)f'(\psi))$ is non-separable, preventing direct integration of Eq. (20). In future work we will fold interdependency of $j_\phi, p'$ and $f(\psi)f'(\psi)$ together to enable calculation of $P(j_\phi(\psi), p'(\psi), f(\psi)f'(\psi))$.

### 3.1 Inference of toroidal flux function

Some progress can be made if we assume the plasma obeys ideal force balance, and $p'(\psi)$ is assumed to be independent of $f(\psi)f'(\psi)$ and $j_\phi(\psi)$. In this instance, and providing we are able to estimate $p'(\psi)$ and its distribution from measurements, then $f(\psi)f'(\psi)$ across the midplane can be computed through Eq. (17). As a first step to inference of force balance in a real plasma, we compute $f(\psi)f'(\psi)$, integrate to find $f(\psi)$, and substitute this back into MINERVA. By examining the changes in the position of the magnetic axis and $q$ profile from the recomputed solution, we are able to quantify the impact of poloidal currents on current tomography without the need for a separate magnetic reconstruction by EFIT. This folds $p'(\psi)$ into $j_\phi(\psi)$ through the influence of poloidal currents.

For an ideal gas, the pressure is given by

$$p = n_e k_B T_e + n_i k_B T_i$$

with $n_e, n_i$ the electron and ion density, and $T_e, T_i$ the electron and ion thermal temperature. At 320 ms, high resolution Thomson scattering measurements of $T_e$ and $n_e$ are available. The data is distributed normally about each data point, and a mean and standard deviation provided. To obtain a pressure estimate, we have assumed that the mean values of density and temperature satisfy $n_i/n_e \approx 0.8$, and $T_i/T_e = 1.1$, as typically extracted from a charge exchange recombination measurement. We have also assumed the same distribution of data as Thomson Scattering.

We have next computed a fit for the pressure $p(\psi)$ using Monte Carlo simulation [22]. The procedure is as follows: at each radial grid point, the inverse transformation method is used to generate $n_e, T_e, n_i, T_i$ samples that satisfying the prescribed pdfs $P(n_e), P(T_e), P(n_i), P(T_i)$. To compute $p(\psi)$ and its distribution we have mapped each sample $P(r)$ to $p(\psi)$ using $\phi(r)$ across the inboard chord determined from MSE. To find a smooth $p(\psi)$ we have then fitted a fourth order polynomial in normalized flux. The sampling process is repeated until the pdf for $p(\psi)$ no longer changes.

Figure 4 shows the sampled pressure and fit for the pressure using $N_s = 2000$ samples. Overlaid is the polynomial fit for each sample across the midplane. The width of the polynomial fit (shown in black) yields a lower case estimate to the standard deviation in $p(\psi)$.
poloidal plasma currents are paramagnetic, and increase is in progress to investigate the impact of the correction to the geometry of flux surfaces or the Shafranov shift. Work change in poloidal flux across the plasma, but not change should lift the on-axis safety factor ≈ at the edge seen in panel (c).

Next, we have integrated $f(\psi) f'(\psi)$ to obtain $f(\psi)$ from panel (c). This yields a toroidal flux profile that varies approximately linearly in poloidal flux from $f(\psi) = -0.407$ at the edge to $f(\psi) = -0.496$ at the core. That is, poloidal plasma currents are paramagnetic, and increase $B_i$ by 22% at the core. Qualitatively, this poloidal current should lift the on-axis safety factor ≈ 22%, increase the change in poloidal flux across the plasma, but not change the geometry of flux surfaces or the Shafranov shift. Work is in progress to investigate the impact of the correction to $f(\psi)$ in MINERVA.

4. Inference of mode structure

A second objective of the project is to test predictions of MHD mode theory, particularly mode structure. To achieve this goal we intend to develop a Bayesian inference tool to identify oscillations using the mode structure of a collection of candidate modes. The low beta heli- liac plasma configuration of H-1NF is ideally suited to this task, as there is little uncertainty in the equilibrium, and precision control of the helical coil currents enables access to a wide range of magnetic rotational transform ($\iota$) profiles, as shown in Fig. 6. The combination thus offers a rich testing ground for Bayesian mode analysis.

For inference of mode structure, the relevant plasma parameters $I$ are the set of eigenmode-specifying parameters (frequency, mode numbers, radial structure moments) and $D$ is a set of Minnov array measurements made outside the plasma. By solving the eigenmode problem in the plasma, and computing the vacuum-region magnetic field corresponding to a given plasma mode we obtain the forward model for $D$ given $I$. Given some prior $P(I)$, the likelihood $P(D|I)$ can then be constructed.

In the remainder of this section we derive the forward model for Global Alfvén Eigenmodes (GAEs) in H-1 plasmas. As a starting point to this sub-task, we compute GAEs in a cylindrical plasma model (coordinates $(r, \theta, z)$) incorporating a vacuum region in the region $r_p < r < r_w$ which is encased by a perfectly conducting wall at $r = r_w$.

GAEs are discrete modes that accumulate at the minimum $\omega^2_{\text{min}}$ of the Alfvén continuum $\omega^2_A = k^2_\parallel \nu^2_A$[23] where $k_\parallel$ denotes the component of the wave vector parallel to the equilibrium magnetic field and $\nu_A$ denotes the Alfvén speed. Expanding $k_\parallel = k \cdot B$ and substituting into $\omega^2_A = k^2_\parallel \nu^2_A$ yields $\omega^2_A \approx (n - m)^2/\rho$ where $m, n$ are the poloidal and toroidal mode numbers respectively, and $\rho$ is the mass density profile, assumed to take the form $\rho = \rho_0 \left(1 - r^2/r_p^2 \right)$ with $\rho_0$ the density on-axis. We define $\omega^2_{\text{min}}$ to be the global minimum of $\omega^2_A$ over the interval $(0, r_p)$, if such a minimum exists. That is, $\omega^2_{\text{min}}$ is the smallest $\omega^2_A(r_0)$ with both $\omega^2_A(r_0) = 0$ and $\omega^2_A(r_0) > 0$ for some radius $r_0$ such that $0 < r_0 < r_p$. Figure 6 shows the dependency of $\omega^2_{\text{min}}$ on $k_\parallel$ for different $m, n$ modes. We have restricted our attention to the frequency interval $0 < f < 40$ kHz, where $f = \omega/(2\pi), \omega$ is the frequency of GAEs, $\omega$ is the Fourier component of the GAEs in an increasing number of radial nodes the closer the eigenfrequency lies to $\omega^2_{\text{min}}$.

We have modelled the plasma region by a stellarator normal-mode formulation in a cylindrical plasma [24, 25], corrected for non constant density profile. The eigenvalue equation for this problem can be written

$$L \varphi = \omega^2 M \varphi,$$

where $\omega$ is given in units of inverse Alfvén time $\tau_A^{-1} = B_0/ R \sqrt{\mu_0 I_p}$, with $B_0$ the field strength at the magnetic axis,
Figure 7: A plot of the minimum in the continuum $\omega_{\text{min}}/(2\pi)$ vs \(\kappa_h\) for a selection of mode numbers \((m,n)\). The cross-hairs mark the location of the eigenmode shown in Fig. 8.

and where \(\varphi\) is defined in terms of the radial element of the fluid displacement

$$r\xi_r = \exp\left[i(m\theta - \frac{n}{R}z)\right]\varphi(r),$$  \hspace{1cm} (23)

with \(i = \sqrt{-1}\). The operators \(L\) and \(M\) are given by:

$$L = -\frac{1}{r} \frac{d}{dr} (n-m) r \frac{d}{dr} + \frac{m^2}{r^2} \left[(n-m)^2 - D_5 + \frac{i}{m} (n-m)\right],$$  \hspace{1cm} (24)

$$M = -\frac{1}{r} \frac{d}{dr} \frac{\rho_r}{\rho_0} r \frac{d}{dr} + \frac{\rho_r (n-m)^2}{\rho_0 r^2}.$$  \hspace{1cm} (25)

Equation (22) is derived by averaging over helical ripple and assuming large aspect ratio. The Suydam stability parameter is given by

$$D_5 = -\frac{\beta_0}{2\epsilon^2} p'(r)\Omega'(r)$$  \hspace{1cm} (26)

where the average field line curvature is

$$\Omega(r) = \epsilon^2 N \left( r^2 \hat{\kappa} + 2 \int r \, \rho \, dr \right)$$  \hspace{1cm} (27)

Here \(j = \sqrt{-1}\), ‘prime’ denotes radial derivatives, \(\hat{\kappa} = r \hat{\kappa}'\), \(p(r)\) is the normalized pressure, \(\epsilon\) is the inverse aspect ratio, \(N\) is the number of turns made by the helical windings and \(\beta_0 = 2\mu_0 p_0 / B_0^2\) is the ratio of plasma pressure to magnetic pressure at magnetic axis.

Turning now to the vacuum region, the magnetic field \(B\) must satisfy \(\nabla \times B = 0\) and \(\nabla \cdot B = 0\). In cylindrical geometry we can write \(B\) in terms of modified Bessel functions, up to a scale factor.[26] The boundary conditions

$$[[B_r]] = [[B_\theta]] = 0$$  \hspace{1cm} (28)

where \([[B]] = B_{\text{vacuum}} - B_{\text{plasma}}\), together with

$$\xi_r(0) = 0, \ |m| \neq 1,$$  \hspace{1cm} (29)

$$\xi_r'(0) = 0, \ |m| = 1,$$  \hspace{1cm} (30)

$$B_r(r_u) = 0,$$  \hspace{1cm} (31)

relate the vacuum solution to that in the plasma.

If \(\omega_{\text{min}}\) appears in the plasma and is positive, a shooting method is used to find solutions to Eqs. (22), (28), (31), and Eq. (29) or Eq. (30) with frequency \(0 < \omega < \omega_{\text{min}}\). More explicitly, we select the appropriate boundary condition for \(\xi_r\) at \(r = 0\), and then ‘shoot’ out to the plasma/vacuum interface, searching frequency space in the region \(0 < \omega < \omega_{\text{min}}\) for eigenfrequencies that are consistent with Eq. (28). Figure 8 shows an example of an \((m,n) = (3,4)\) eigenmode found in this manner, located at the point \((\kappa_h, \omega)\) indicated by the cross-hairs in Figure 7. The global structure seen has promising features when compared to recent spectral measurements of the mode: the mode structure is global in radial extent, and the radial position of the peaks nodes broadly matches observations.

Figure 8: The radial fluid element displacement \(\xi_r\) as a function of minor radius for \((m,n) = (3,4)\) and \(\kappa_h = 0.73\) at the eigenfrequency indicated. This eigenmode has angular frequency \(\omega = 12.1 \, \text{k rad s}^{-1}\).

5. Conclusions

In this work we have described a Bayesian inversion framework for inference of force-balance and mode structure. Progress has been reported in three areas: development of Bayesian inversion for current profiles; inference of the toroidal flux profile assuming the plasma satisfied ideal MHD; and calculation of a Global Alfvén eigenmodes in a helical-ripple averaged cylindrical plasma model.

Based on working by Svensson et al. we have developed Bayesian inversion for current profiles and poloidal flux contours in MAST. The model comprises a grid of rectangular toroidally extended current beams, and is constrained by pickup coils, flux loops and MSE data. Owing to the placement of the MSE viewing optics the tangent of the polarisation angle of emitted light is linear in \(B_R\) and \(B_Z\). Signals from the pickup coils and flux loops are also linear in \(B_R\) and \(B_Z\). This feature means the prediction vector for the data is linear in the currents. We have used a conditional autoregressive prior to describe the initial current distribution, which has the advantage of enforcing some smoothness between adjacent current beams while minimizing spatially long ranged effects. Posterior calculations of the poloidal flux contours of MAST #22254 at 320 ms illustrate good boundary agreement to EFIT.
In Sec. 3 we outlined an inference technique for the extraction of force balance, as applied to the Grad Shafranov equation. Using Thomson scattering data and a Monte Carlo simulation technique, we fitted pressure profiles to the inboard midplane chord of MAST, and computed \( p'(\psi) \). Assuming ideal MHD force balance we then calculated \( f(\psi)f'(\psi) \). While the \( f(\psi)f'(\psi) \) profile is similar to EFIT, the change in \( f(\psi) \) from the core to the edge is 22\%, which is slightly larger than the EFIT value. Qualitatively similar to EFIT, the change in \( \kappa \) is a function of helical winding current ratio for different plasma, and the minimum in the continuum computed as \( \psi \) increases. Hence, we determined at an angular frequency of 12.1 krad s\(^{-1}\) we intend to transform the cylindrical eigenfunction to a poloidal eigenfunction to validate this work. The Bayesian model has been applied to the range of \( j_\phi \) profiles accessible to H-1 plasmas, and the minimum in the continuum computed as a function of helical winding current ratio \( \kappa_\phi \). The eigenfunction of a candidate \((m,n) = (3,4)\) GAE mode was determined at an angular frequency of 12.1 krad s\(^{-1}\). The eigenfunction has similar structure to recent observations in H1, including the same number of radial nodes which occur at similar radial positions.

In ongoing work we are developing forward models of Thomson scattering and charge exchange recombination spectroscopy for inference of force balance in MAST plasmas. We will be exploiting this information not only to improve equilibrium reconstruction, but to use Bayesian inference as a tool with which to resolve competing equilibrium models. In particular, we intend to compare force balance descriptions of ideal MHD, ideal MHD with flow [11], two fluid models [27], and energetic fluid resolved equilibria [2]. The Bayesian model offers a rigorous framework which can quantify the fit and thereby elucidate the important underlying physics. With regards mode structure, we intend to transform the cylindrical eigenfunction to a beam cross-section equilibria [28] and identify a candidate mode set which spans the different type of mode structures observed. The next stage of this work involves deployment of forward models for mode structures into MINERVA to compute the posterior for the mode class and parameters. On a longer time scale we envisage the cylindrical plasma model will be replaced by a fully 3D ideal MHD wave mode code, CAS3D [29]. As with equilibrium modelling, the importance of mode analysis by Bayesian inversion is that it offers a rigorous framework in which to identify wave mode structures, some of which have deleterious effects on plasma performance.

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